# Analytical Solution with Two Time Scales of Circular Restricted Three-Body Problem 

F.B. Gao ${ }^{1}$, H.L. $\mathrm{Hu}^{2}$<br>${ }^{1}$ School of Mathematical Science, Yangzhou University, Yangzhou 225002, China<br>${ }^{2}$ College of Mathematics, Physics and Information Engineering, Zhejiang Normal University, Jinhua 321004, China


#### Abstract

Analytical solution performs a vital role in a wide variety of deep space exploration missions, especially the periodic solutions which were believed to be the unique avenue to solve three-body problem by Poincaré. As the absence of a general solution for the problem, an approximate analytical solution of the circular restricted three-body problem is addressed by employing multiple scales method in conjunction with some analytical techniques. It is worthwhile to note that the presented solution in three-dimensional space contains two time scales: $t$ and $\varepsilon t$ ( $\varepsilon$ is a small, dimensionless parameter), which is significative to improve and perfect the known literature.


Keywords-Analytical solution, multiple scales, three-body problem.

## I. INTRODUCTION

An existence of conditional periodic solutions of circular restricted three-body problem (CR3BP) was presented by Gao [1]. However, because of lacking exact analytical solutions for the problem, a considerable amount of researchers were attracted to develop the approximate analytical solutions of the problem. In 1973, Farquhar and Kamel [2] developed an approximation method for this type of orbit. Richardson [3, 4] presented a third-order analytical solution about the collinear libration points of the CR3BP based on the method of successive approximations and a technique similar to the LindstedtPoincaré method. Lu and Zhao [5] put forward a kind of improved third-order approximate analytical solution in 2009, which is more accurate than the classical analytic solution of Richardson. In addition, Nayfeh [6, 7] studied two types of resonance near the planar triangular libration points under a case of small amplitude. The planar 3:1 resonance with time scales $t$ and $\varepsilon^{2} t$ was discussed when the planar restricted three-body problem was expanded to second-order terms in $\mathcal{E}$. Moreover, approximate analytical solutions of the planar $2: 1$ resonance with time scales $t$ and $\varepsilon t$ was also investigated when the problem was expanded to third-order terms in $\varepsilon$. These creative results provide us with new motivation in studying the solutions of the three-body problem.

In this paper, multiple scales method (see Nayfeh [8]) will be employed to construct a three-dimensional (3D) approximate analytical solution of a spatial CR3BP. The addressed 3D solution will be written in the following form

$$
\begin{equation*}
\mathbf{u}=\varepsilon \mathbf{u}_{1}(t, \varepsilon t)+\varepsilon^{2} \mathbf{u}_{2}(t, \varepsilon t) \tag{1}
\end{equation*}
$$

where $\mathbf{u}=(x, y, z), \mathbf{u}_{i}=\left(x_{i}, y_{i}, z_{i}\right), i=1,2$, and $\varepsilon$ is a small, dimensionless parameter.
It is worthwhile to note that when the problem is expanded to third-order, the above presented solution in 3D space contains two time scales, which is significative to improve and perfect the known literature, in which the constructed solutions mainly possess three cases: a) The three-body problem discussed is a planar case, so the proposed approximate solution is a twodimensional curve. b) The 3D approximate solution contains only one time scale when the spatial CR3BP was expanded to the third-order. c) The 3D approximate solution contains two time scales when the spatial CR3BP was expanded to the second-order.

## II. CONSTRUCTION OF THE 3D APPROXIMATE ANALYTICAL SOLUTION

Note that the third-order approximate system with small amplitude of the CR3BP in rotating frame can be described as follows (the general third-order approximate system can be found at Koon et al. [9])

$$
\begin{equation*}
\ddot{x}-2 \dot{y}-\left(1+2 c_{2}\right) x=\frac{3}{2} \varepsilon c_{3}\left(2 x^{2}-y^{2}-z^{2}\right)+2 \varepsilon c_{4}\left(2 x^{2}-3 y^{2}-3 z^{2}\right), \tag{2a}
\end{equation*}
$$

$$
\begin{align*}
& \ddot{y}+2 \dot{x}+\left(c_{2}-1\right) y=-3 \varepsilon c_{3} x y-\frac{3}{2} \varepsilon c_{4} y\left(4 x^{2}-y^{2}-z^{2}\right),  \tag{2b}\\
& \ddot{z}+c_{2} z=-3 \varepsilon c_{3} x z-\frac{3}{2} \varepsilon c_{4} z\left(4 x^{2}-y^{2}-z^{2}\right), \tag{2c}
\end{align*}
$$

where

1. $c_{n}=\frac{1}{\gamma_{L}^{3}}\left[( \pm 1)^{n} \mu+(-1)^{n} \frac{(1-\mu) \gamma_{L}^{n+1}}{\left(1 \mp \gamma_{L}\right)^{n+1}}\right]$, the upper sign is for $L_{1}$ and the lower one for $L_{2}$,
$c_{n}=\frac{1}{\gamma_{L}^{3}}\left[1-\mu+\frac{\mu \gamma_{L}^{n+1}}{\left(1+\gamma_{L}\right)^{n+1}}\right]$, for libration point $L_{3}$,
$\gamma_{L}=n_{1}^{\frac{2}{3}}, n_{1}$ is the angular velocity of relative movement between two primaries.
The solution of equations (2a) $\sim(2 c)$ are assumed to possess the following form

$$
\begin{align*}
& x=\varepsilon x_{1}\left(T_{0}, T_{1}\right)+\varepsilon^{2} x_{2}\left(T_{0}, T_{1}\right),  \tag{3a}\\
& y=\varepsilon y_{1}\left(T_{0}, T_{1}\right)+\varepsilon^{2} y_{2}\left(T_{0}, T_{1}\right),  \tag{3b}\\
& z=\varepsilon z_{1}\left(T_{0}, T_{1}\right)+\varepsilon^{2} z_{2}\left(T_{0}, T_{1}\right), \tag{3c}
\end{align*}
$$

where $T_{0}=t, T_{1}=\varepsilon t$.

Then, time derivatives become $\frac{d}{d t}=D_{0}+\varepsilon D_{1}+\cdots, \frac{d^{2}}{d t^{2}}=D_{0}^{2}+2 \varepsilon D_{0} D_{1}+\cdots, D_{k}=\frac{\partial}{\partial T_{k}}, k=0,1, \cdots$.

Substituting equations (3) into equations (2a) $\sim(2 \mathrm{c})$ and equating the coefficients of $\varepsilon, \varepsilon^{2}$, and $\varepsilon^{3}$ to zero, leads to

$$
\begin{align*}
& D_{0}^{2} x_{1}-2 D_{0} y_{1}-\left(1+2 c_{2}\right) x_{1}=0  \tag{4a}\\
& D_{0}^{2} y_{1}+2 D_{0} x_{1}+\left(c_{2}-1\right) y_{1}=0  \tag{4b}\\
& D_{0}^{2} z_{1}+c_{2} z_{1}=0  \tag{4c}\\
& D_{0}^{2} x_{2}+2 D_{0} D_{1} x_{1}-2 D_{0} y_{2}-2 D_{1} y_{1}-\left(1+2 c_{2}\right) x_{2}=0  \tag{5a}\\
& D_{0}^{2} y_{2}+2 D_{0} D_{1} y_{1}+2 D_{0} x_{2}+2 D_{1} x_{1}+\left(c_{2}-1\right) y_{2}=0,  \tag{5b}\\
& D_{0}^{2} z_{2}+2 D_{0} D_{1} z_{1}+c_{2} z_{2}=0  \tag{5c}\\
& 2 D_{0} D_{1} x_{2}-2 D_{1} y_{2}=\frac{3}{2} c_{3}\left(2 x_{1}^{2}-y_{1}^{2}-z_{1}^{2}\right), \tag{6a}
\end{align*}
$$

$$
\begin{align*}
& 2 D_{0} D_{1} y_{2}+2 D_{1} x_{2}=-3 c_{3} x_{1} y_{1}  \tag{6b}\\
& 2 D_{0} D_{1} z_{2}=-3 c_{3} x_{1} z_{1} \tag{6c}
\end{align*}
$$

Let secular terms be equal to zero, the general solution of equations (4a) $\sim(4 \mathrm{c})$ and $(5 \mathrm{a}) \sim(5 \mathrm{c})$ read

$$
\begin{align*}
& x_{1}=\rho_{1} \cos (\lambda t)  \tag{7a}\\
& y_{1}=-\delta \rho_{1} \sin (\lambda t)  \tag{7b}\\
& z_{1}=\rho_{2} \cos (\lambda t)  \tag{7c}\\
& x_{2}=\alpha_{1}\left(T_{1}\right) \cos \left(\lambda t+\beta_{1}\left(T_{1}\right)\right)  \tag{8a}\\
& y_{2}=-\delta \alpha_{1}\left(T_{1}\right) \sin \left(\lambda t+\beta_{1}\left(T_{1}\right)\right)  \tag{8b}\\
& z_{2}=\alpha_{2}\left(T_{1}\right) \cos \left(\lambda t+\beta_{2}\left(T_{1}\right)\right) \tag{8c}
\end{align*}
$$

where $\rho_{1}, \rho_{2}$ and $\delta$ are constants, $\lambda$ is the mould of the pure imaginary roots of equations (4), $\alpha_{i}$ and $\beta_{i}(i=1,2$ ) satisfy

$$
\begin{align*}
& \alpha_{1}^{\prime}\left(T_{1}\right)= \frac{3}{8(\delta-\lambda)} c_{3}\left[\left(\left(2+\delta^{2}\right) \rho_{1}^{2}-\rho_{2}^{2}\right) \cos (2 \lambda t)+\left(2-\delta^{2}\right) \rho_{1}^{2}-\rho_{2}^{2}\right] \sin \left(\lambda t+\beta_{1}\left(T_{1}\right)\right) \\
&+\frac{3}{4(1-\lambda \delta)} c_{3} \delta \rho_{1}^{2} \sin (2 \lambda t) \cos \left(\lambda t+\beta_{1}\left(T_{1}\right)\right),  \tag{9a}\\
& \alpha_{1} \beta_{1}^{\prime}\left(T_{1}\right)= \frac{3}{8(\delta-\lambda)} c_{3}\left[\left(\left(2+\delta^{2}\right) \rho_{1}^{2}-\rho_{2}^{2}\right) \cos (2 \lambda t)+\left(2-\delta^{2}\right) \rho_{1}^{2}-\rho_{2}^{2}\right] \cos \left(\lambda t+\beta_{1}\left(T_{1}\right)\right) \\
&-\frac{3}{4(1-\lambda \delta)} c_{3} \delta \rho_{1}^{2} \sin (2 \lambda t) \sin \left(\lambda t+\beta_{1}\left(T_{1}\right)\right),  \tag{9b}\\
&-2 \lambda\left[\alpha_{2}^{\prime} \sin \left(\lambda T_{0}+\beta_{2}\right)+\alpha_{2} \beta_{2}^{\prime} \cos \left(\lambda T_{0}+\beta_{2}\right)\right]=-\frac{3}{2} c_{3} \delta \rho_{1} \rho_{2}\left[\cos \left(2 \lambda T_{0}\right)+1\right] . \tag{9c}
\end{align*}
$$

Therefore, a third-order approximate analytical solution with two time scales to equations (2a) $\sim(2 c)$ can be represented as

$$
\begin{align*}
& x=\varepsilon \rho_{1} \cos (\lambda t)+\varepsilon^{2} \alpha_{1}(\varepsilon t) \cos \left(\lambda t+\beta_{1}(\varepsilon t)\right)  \tag{10a}\\
& y=-\varepsilon \delta \rho_{1} \sin (\lambda t)-\varepsilon^{2} \delta \alpha_{1}(\varepsilon t) \sin \left(\lambda t+\beta_{1}(\varepsilon t)\right)  \tag{10b}\\
& z=\varepsilon \rho_{2} \cos (\lambda t)+\varepsilon^{2} \alpha_{2}(\varepsilon t) \cos \left(\lambda t+\beta_{2}(\varepsilon t)\right) \tag{10c}
\end{align*}
$$

where $\alpha_{1}$ and $\beta_{1}$ satisfy equations (9a) and (9b), $\alpha_{2}$ and $\beta_{2}$ satisfy equation (9c).

## III. CONCLUSION

Since the governing equations of the third body is time-dependent in inertial frame, which appear as a high-dimensional nonlinear autonomous system. According to the abundant known literature, this system seems unlikely to be solved via analytical approaches. For example, if we employ multiple scales method, then it is difficult solving the solutions to the equations derived from equating coefficients of $\varepsilon^{2}$ to zero, not to mention the non-autonomous obtained from third-order terms in $\mathcal{E}$. However, these will be possible in rotating frame, where the equations are characterized by nonlinear autonomous system. In this frame, the method of multiple scales in conjunction with some analytical techniques is adopted to construct a 3D approximate analytical periodic solution for CR3BP. The solution was demonstrated with two different time scales $t$ and $\varepsilon t$ when the system was expanded to $\varepsilon^{3}$-order. This result improves and perfects the known literature.

## ACKNOWLEDGEMENTS

The authors gratefully acknowledge the support of the National Natural Science Foundation of China (NNSFC) through grant Nos. 11302187, 11271197 and 61473340; the Ministry of Land and Resources Research of China in the Public Interest through grant No. 201411007.

## REFERENCES

[1] F.B. Gao, W. Zhang, "A study on periodic solutions for the circular restricted three-body problem," Astronomical Journal, vol. 148, 2014, Article ID 116.
[2] R.W. Farquhar, A.A. Kamel, "Quasi-periodic orbits about the translunar libration point," Celestial Mech., vol. 7, 1973, pp. 458-473.
[3] D.L. Richardson, "Halo orbit formulation for the ISEE-3 mission," J. Guid. Control, vol. 3, 1980, pp. 543-548.
[4] D.L. Richardson, "Analytical construction of periodic orbits about the collinear points," Celestial Mech., vol. 22, 1980, pp. $241-253$.
[5] S.T. Lu, Y.S. Zhao, "The improvement of Richardson's three order approximate analytical solution of halo orbit," J. Astr., vol. 30, 2009, pp. 863-869.
[6] A.H. Nayfeh, "Three-to-one resonances near the equilateral libration points," AIAA J., vol. 8, 1970, pp. 2245-2251.
[7] A.H. Nayfeh, "Two-to-one resonances near the equilateral libration points," AIAA J., vol. 9, 1971, pp. 23-27.
[8] A.H. Nayfeh, D.T. Mook, "Nonlinear oscillations," Wiley-Vch Verlag GmbH \& Co. KGaA, 2004.
[9] W.S. Koon, M.W. Lo, J.E. Marsden, S.D. Ross, "Dynamical systems, the three-body problem and space mission design," Marsden Books, ISBN 978-0-615-24095-4, 2011.

